



# SPACESTORM



## Modelling the diffusion due to wave-particle interactions in the radiation belts

Sarah Glauert, Richard Horne, Nigel Meredith  
Toby Kersten

British Antarctic Survey, Cambridge, UK

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# Background

Diffusion equation for phase-space density ( $f$ ) in pitch-angle ( $\alpha$ ), energy ( $E$ ) and  $L^*$  ( $L$ ) coordinates

$$\begin{aligned}
 \frac{\partial f}{\partial t} = & \underbrace{L^2 \frac{\partial}{\partial L} \left( \frac{D_{LL}}{L^2} \frac{\partial f}{\partial L} \right)}_{\text{Radial}} + \frac{1}{g(\alpha)} \frac{\partial}{\partial \alpha} \left( g(\alpha) \left( \underbrace{D_{\alpha\alpha} \frac{\partial f}{\partial \alpha}}_{\text{Pitch-angle}} + \underbrace{D_{\alpha E} \frac{\partial f}{\partial E}}_{\text{Mixed pitch-angle/energy}} \right) \right)_{\alpha, EL} \\
 & + \frac{1}{A(E)} \frac{\partial}{\partial E} \left( A(E) \left( \underbrace{D_{E\alpha} \frac{\partial f}{\partial \alpha}}_{\text{Mixed pitch-angle/energy}} + \underbrace{D_{EE} \frac{\partial f}{\partial E}}_{\text{Energy}} \right) \right)_{\alpha, EL} - \underbrace{\frac{f}{\tau(\alpha, E, L)}}_{\text{Loss processes}}
 \end{aligned}$$

$$A(E) = (E + E_0)(E + 2E_0)^{\frac{1}{2}} E^{\frac{1}{2}}$$

$$g(\alpha) = \sin \alpha \cos \alpha (1.30 - 0.56 \sin \alpha)$$

# Background

BAS-RBM solves a diffusion equation for phase-space density ( $f$ ) in pitch-angle ( $\alpha$ ), energy ( $E$ ) and  $L^*$  ( $L$ ) coordinates

$$\frac{\partial f}{\partial t} = L^2 \frac{\partial}{\partial L} \left( \frac{D_{LL}}{L^2} \frac{\partial f}{\partial L} \right) \Bigg|_{\mu J} + \frac{1}{g(\alpha)} \frac{\partial}{\partial \alpha} \left( g(\alpha) \left( D_{\alpha\alpha} \frac{\partial f}{\partial \alpha} + D_{\alpha E} \frac{\partial f}{\partial E} \right) \right) \Bigg|_{\alpha, EL}$$

$$+ \frac{1}{A(E)} \frac{\partial}{\partial E} \left( A(E) \left( D_{E\alpha} \frac{\partial f}{\partial \alpha} + D_{EE} \frac{\partial f}{\partial E} \right) \right) \Bigg|_{\alpha, EL} - \frac{f}{\tau(\alpha, E, L)}$$

$$A(E) = (E + E_0)(E + 2E_0)^{\frac{1}{2}} E^{\frac{1}{2}}$$

$$g(\alpha) = \sin \alpha \cos \alpha (1.30 - 0.56 \sin \alpha)$$

# How do you calculate the diffusion rates?

Use a code like PADIE [Glauert & Horne, 2005]

- Gyro-resonant wave-particle interactions
- Describe the wave
  - Frequency distribution
  - Wave normal angle distribution
  - Distribution of the waves in space
- Location
  - Plasma frequency
  - Magnetic field
- Particle type (electron or ion) and energy
- Number of resonances



local diffusion coefficients  $D_{\alpha\alpha}$ ,  $D_{\alpha E}$  and  $D_{EE}$

# Drift-averaged diffusion rates

Average the local diffusion coefficients over one bounce

→ Bounce average diffusion coefficients  $\langle D_{\alpha\alpha} \rangle$ ,  $\langle D_{\alpha E} \rangle$  and  $\langle D_{EE} \rangle$

Average around a drift path

→ drift and bounce averaged coefficients used in the model

$$\langle D_{\alpha\alpha} \rangle^d, \langle D_{\alpha E} \rangle^d \text{ and } \langle D_{EE} \rangle^d$$

Repeat for

- multiple energies
- multiple  $L^*$
- varying geomagnetic conditions
  - Plasma frequency
  - Wave parameters

# Chorus diffusion model

Model from EU-FP7 project SPACECAST [Horne et al., 2013]

- Lower ( $0.1 f_{ce} \leq f < 0.5 f_{ce}$ ) and upper  $0.5 f_{ce} < f \leq 0.9 f_{ce}$  band chorus
- Parameterised geomagnetic activity by Kp and AE

- Data from 7 satellites

$f_{ce}$  = local gyro-frequency

- Wave spectra determined for

- $2 \leq L^* \leq 10$  in steps of  $0.5 R_e$
- 5 levels geomagnetic activity
- 3 hour MLT bins
- $6^\circ$  latitude bins  $0 \leq |\lambda| \leq 60^\circ$

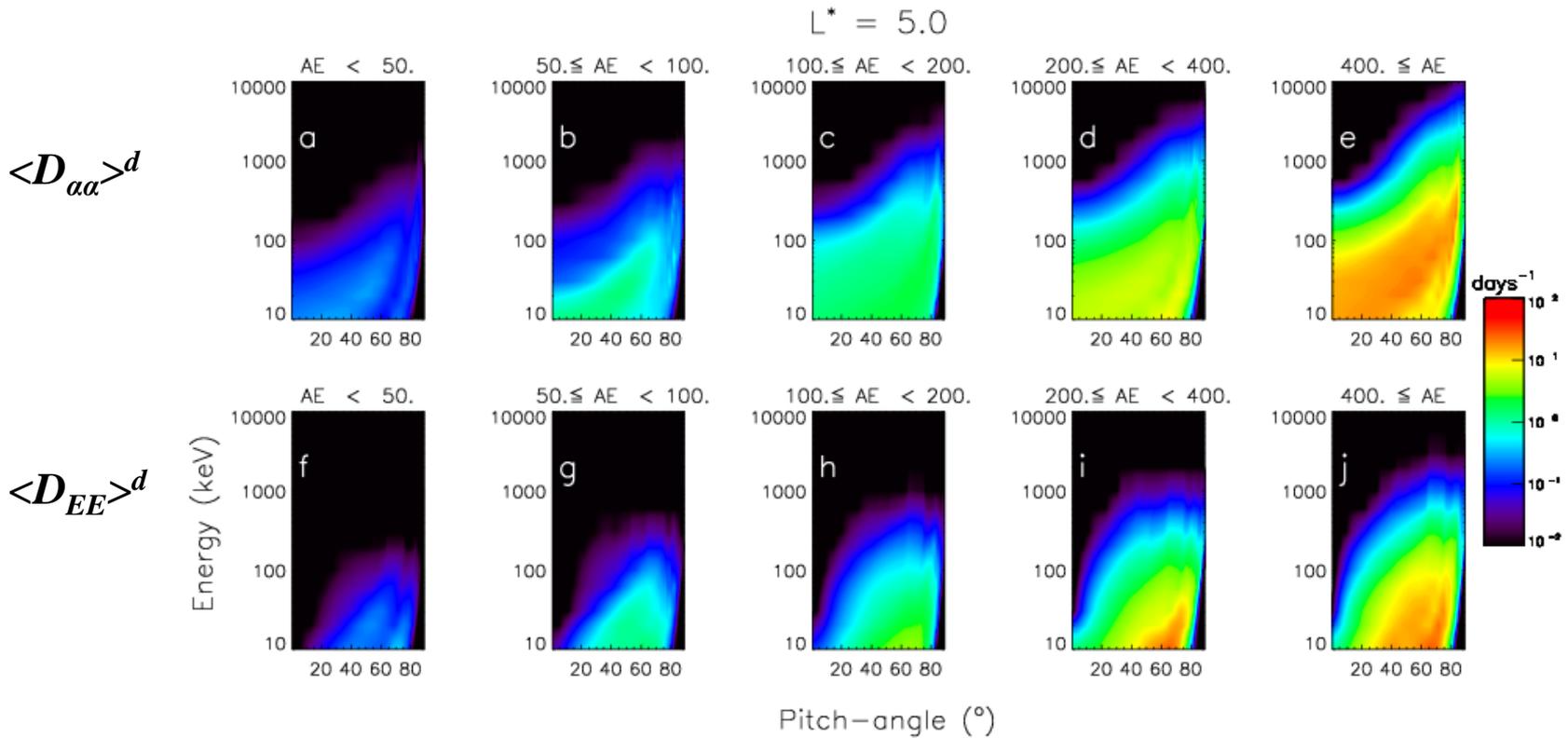
- Wave normal angle distribution - peak at  $0^\circ$ , width of  $\tan 30^\circ$

- $10 \text{ keV} \leq \text{Energy} \leq 30 \text{ MeV}$

- $f_{pe}/f_{ce}$  from new model based on CRRES and THEMIS data

# Chorus diffusion model

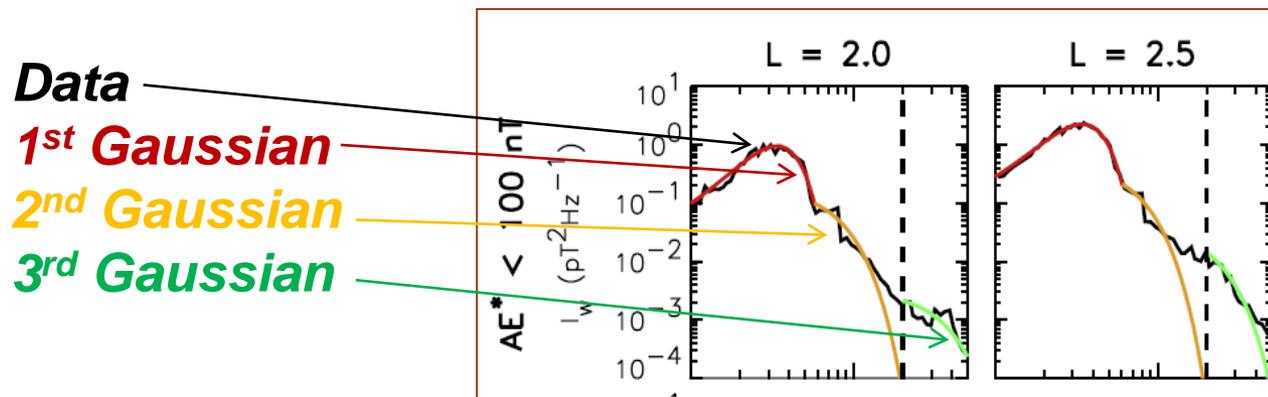
- Diffusion increases with increasing activity
- Significant energy diffusion



# Plasmaspheric hiss diffusion model

Hiss model based on CRRES data [Glauert et al., 2014]

- 100 Hz to 5 kHz – includes lightning generated whistlers
- Wave spectra determined for
  - $2 \leq L^* \leq 6.5$  in steps of  $0.5 R_e$
  - 3 levels geomagnetic activity
  - Average spectra over all MLT for each  $L^*$
  - Fit 3 Gaussian distributions to cover whole frequency range [Meredith et al., 2007]

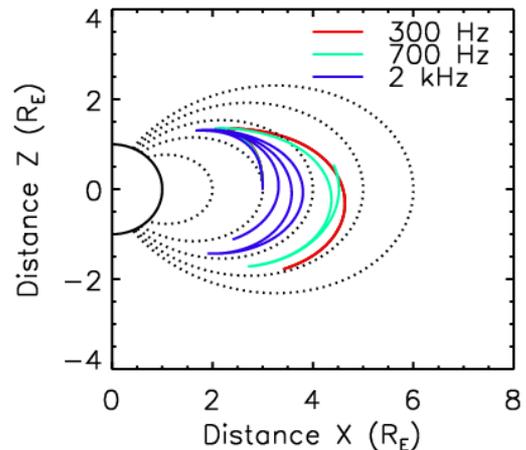


# Wave-normal angle model for hiss

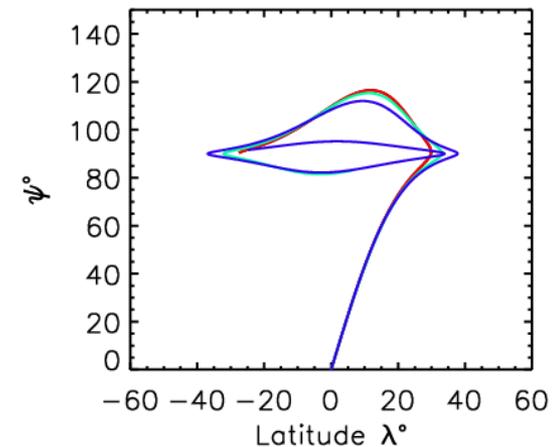
Peak wave-normal angle varies with latitude

- Field aligned near equator, oblique at higher latitudes [*Bortnik et al., 2008*]
- HOTRAY ray tracing code estimates variation with latitude
  - at each L for 3 frequencies, (300 Hz, 700 Hz and 2 kHz)
- Width of the wave-normal angle distribution fixed at  $\tan 80^\circ$

L = 3 Distance Z vs distance X

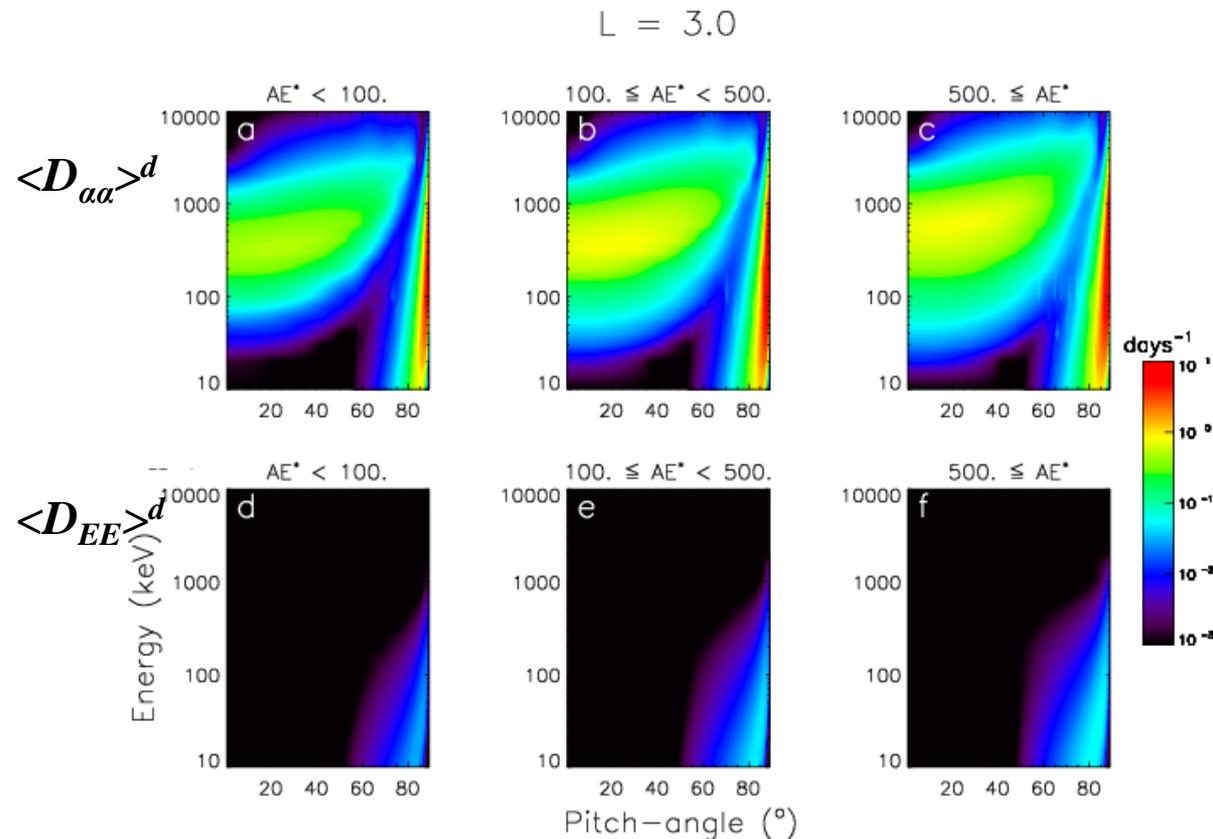


wna vs latitude



# Diffusion due to hiss and LGW

- Losses increase with increasing geomagnetic activity
- Losses peak at ~300 keV at  $L^*=3$
- No significant energy diffusion

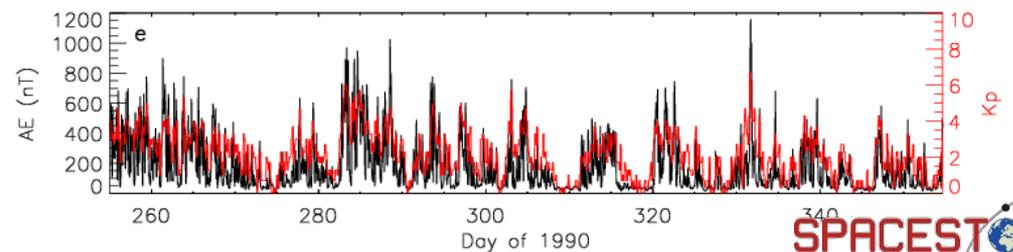
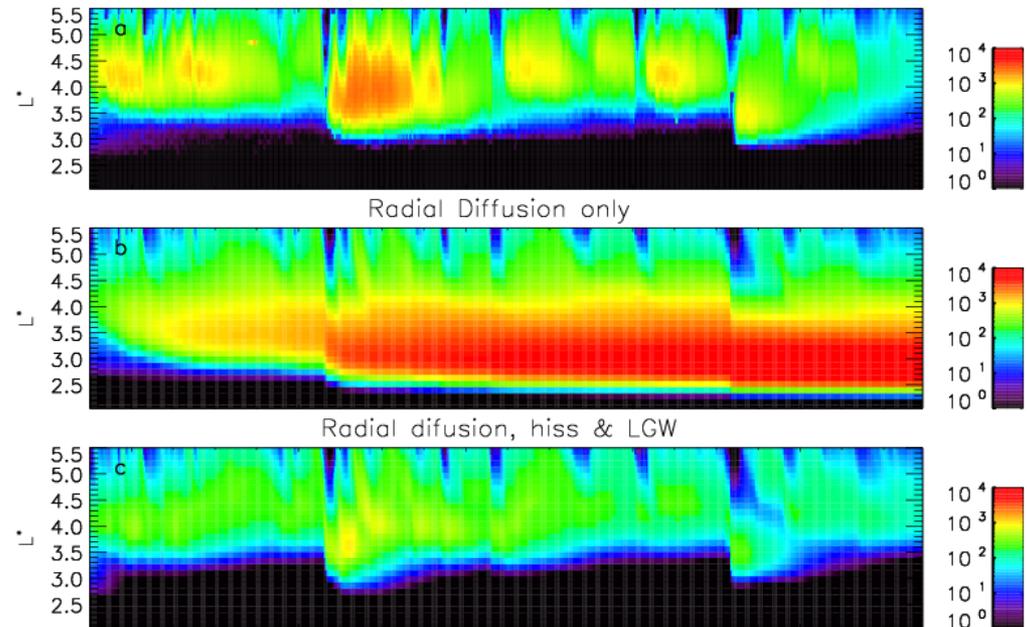


# 3d simulations

- Hiss model reproduces inner edge of outer belt
- Chorus provides acceleration

90° equatorial flux ( $\text{cm}^{-2}\text{sr}^{-1}\text{s}^{-1}\text{keV}^{-1}$ ) for 976 keV electrons

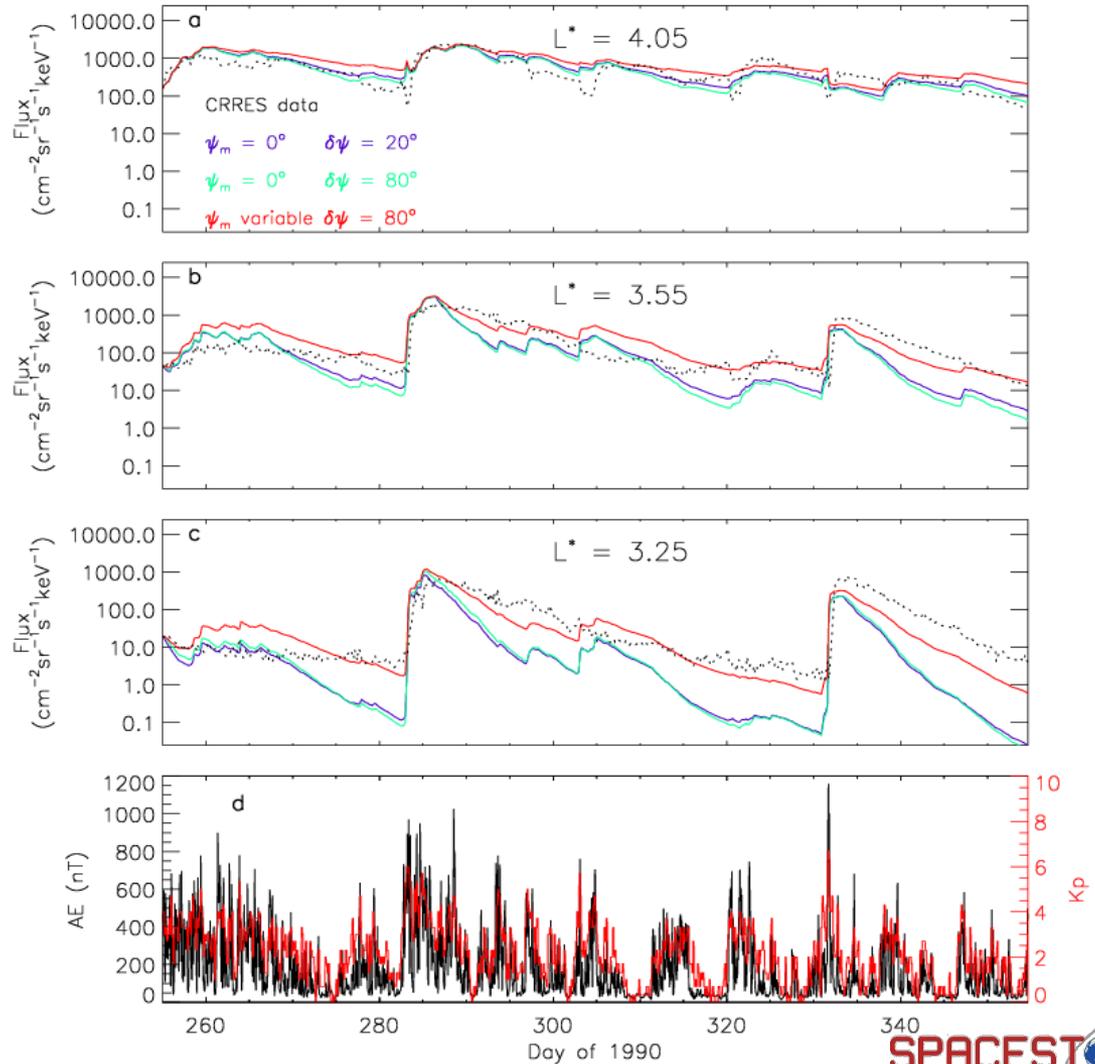
CRRES data



# Decay rates

- Hiss model (red line) reproduces decay rates

90° equatorial flux for 976 keV electrons



Red - Hiss model  
Black - data

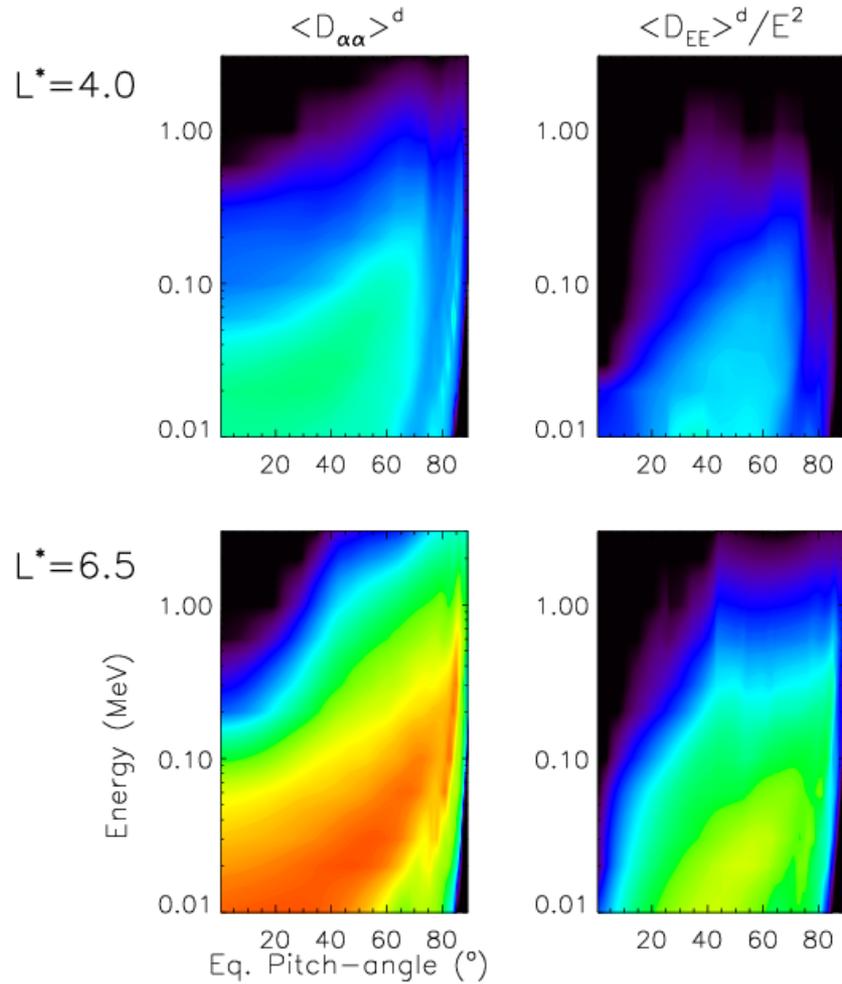
# Improvements to the chorus model

Two recent results:

1. Significant wave power below  $0.1 f_{ce}$  at higher latitudes  
[Meredith et al., 2014]
  - Not included in previous chorus models
2. Wave normal angle distribution width  $\delta X \sim \tan 15^\circ$   
[Santolik et al., 2014]

# Low frequency chorus

## Upper & lower band chorus

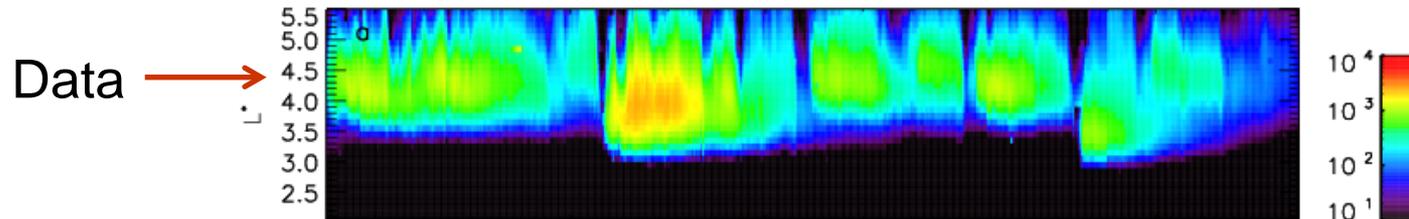


$100 \text{ nT} \leq \text{AE} < 200 \text{ nT}$

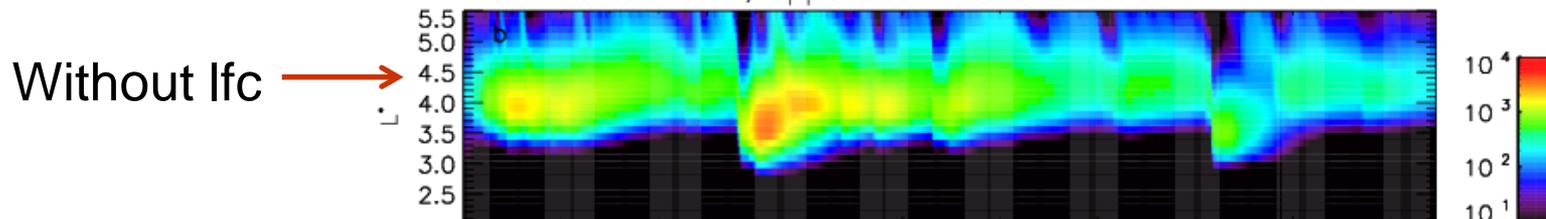
# Simulation with low frequency chorus

90° equatorial flux ( $\text{cm}^{-2}\text{sr}^{-1}\text{s}^{-1}\text{keV}^{-1}$ ) for 976 keV electrons

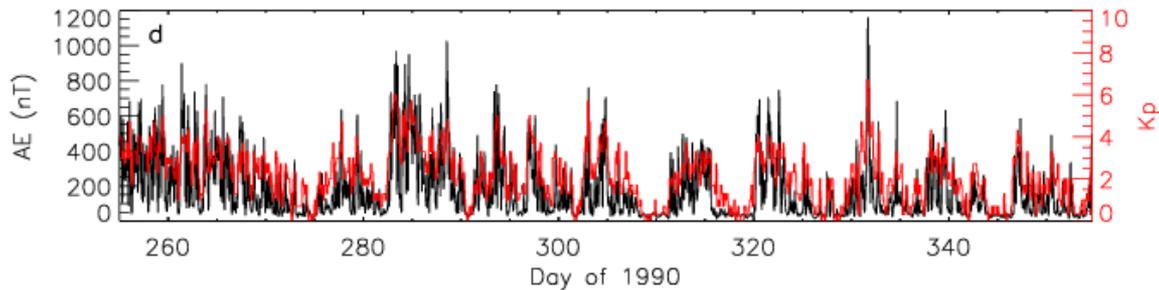
CRRES data



Hiss, upper and lower band chorus



Increased loss



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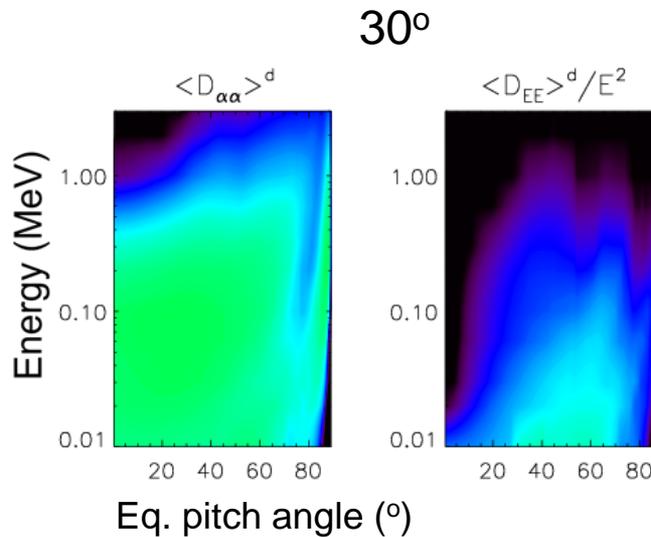
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# Wave normal angle distribution

Recalculated chorus diffusion matrix (again)

- Width -  $\tan 15^\circ$

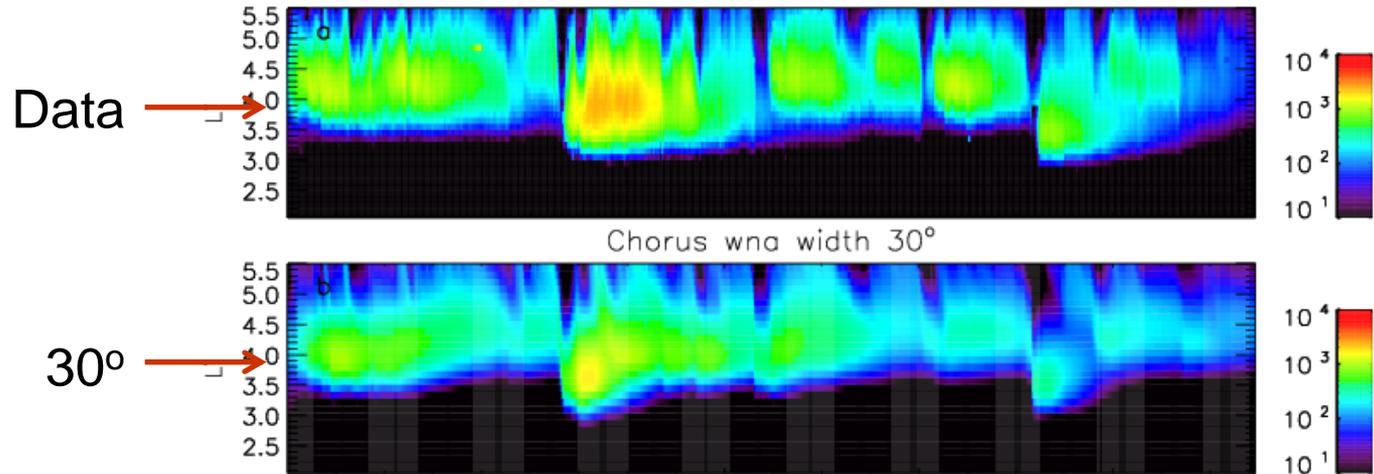
$L^* = 4$   
 $2 \leq Kp < 3$



# Simulation with wna width 15°

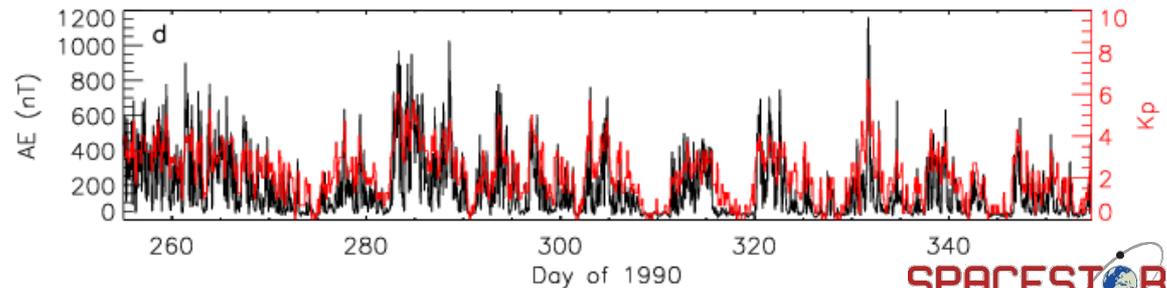
90° equatorial flux ( $\text{cm}^{-2}\text{sr}^{-1}\text{s}^{-1}\text{keV}^{-1}$ ) for 976 keV electrons

CRRES data



Small increase in acceleration

Increase is more significant at higher energies  $\sim 4$  MeV



# Improvements to the hiss model

- Have to include a plasmopause model
- Have one spectra for all MLT, scaled by wave power
- Agapitov et al. [2013] suggest  $\delta X \sim \tan 20^\circ$ 
  - Our model has  $\delta X \sim \tan 80^\circ$

Toby is working on a new model ...

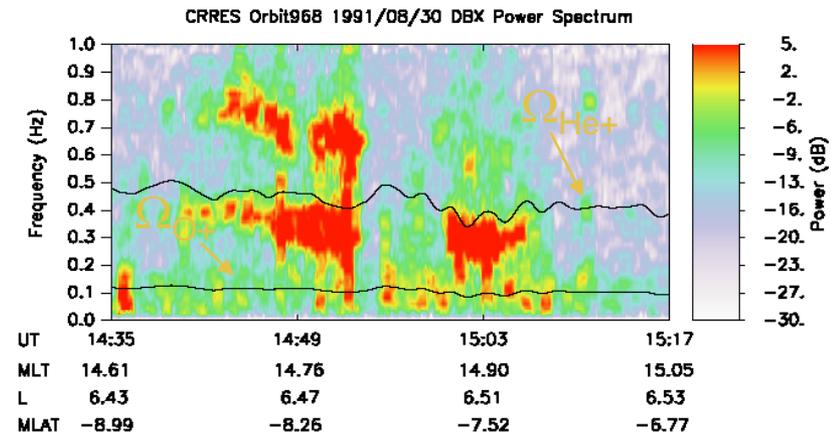
- Data from 7 satellites
- New method of distinguishing hiss from chorus
- 1 hour resolution in MLT
- Calculate wave spectra for each MLT and  $L^*$
- Update wave-normal angle model
- 6 levels of geomagnetic activity



# Electromagnetic ion cyclotron (EMIC) waves

## Diffusion matrix for EMIC waves

- SPACECAST and MAARBLE projects [Kersten et al., 2014]
- Averaged properties of the waves
- CRRES data
- Hydrogen and helium band waves





# Effect of EMIC waves

Same 100 days

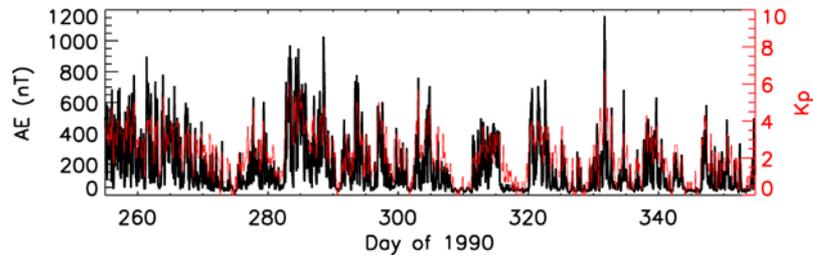
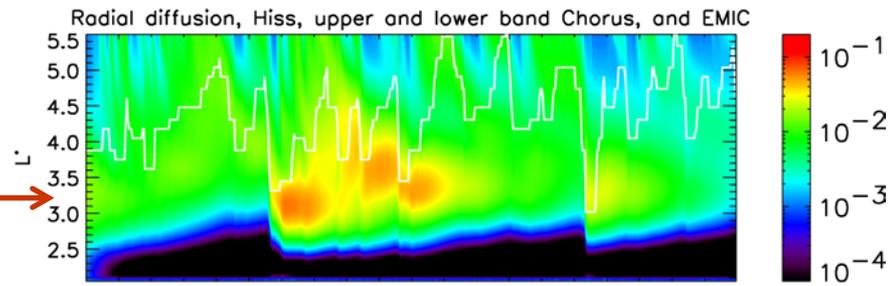
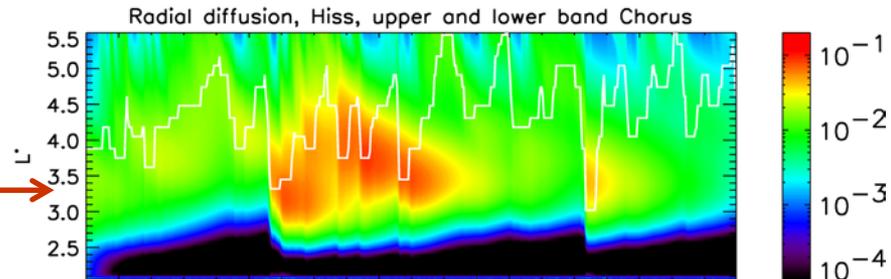
45°, 6 MeV electrons

No EMIC waves

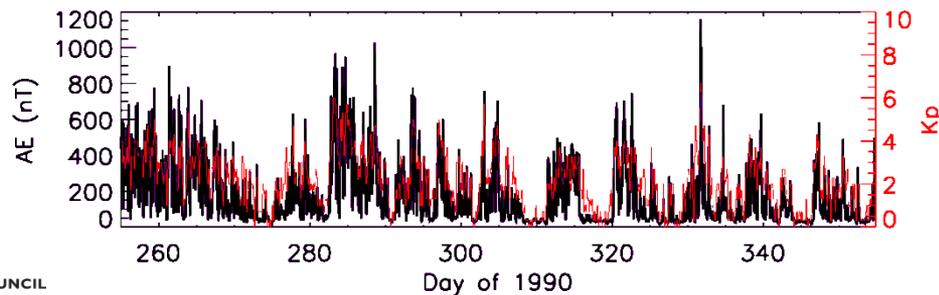
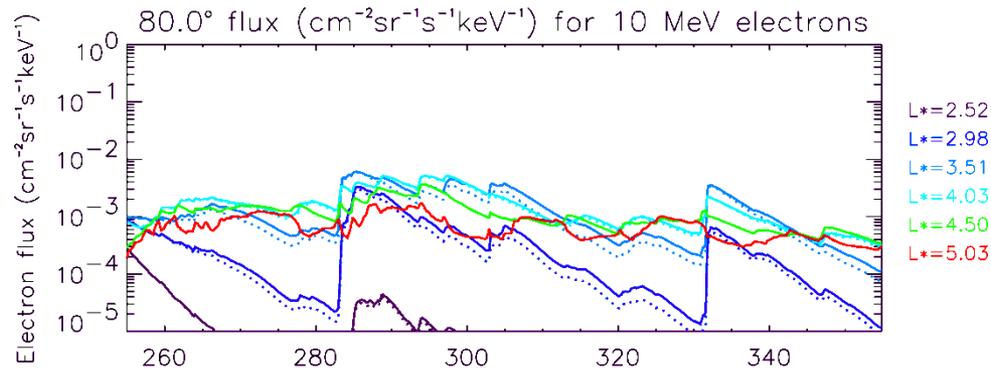
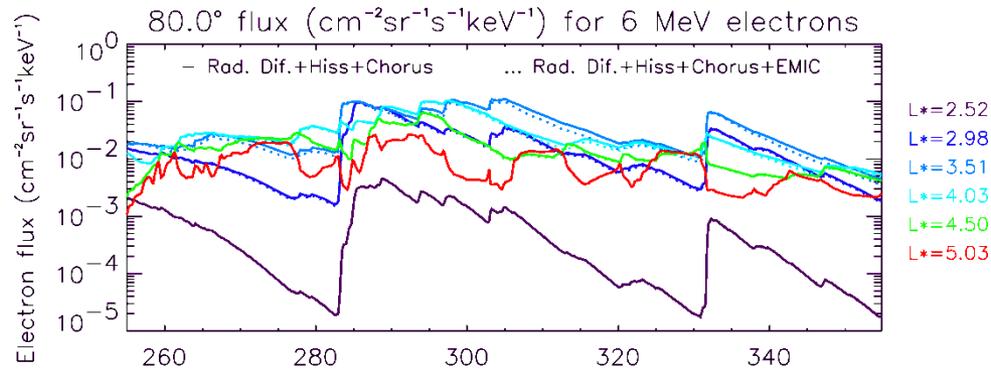
Including EMIC waves

- Increased losses

45.0° flux ( $\text{cm}^{-2}\text{sr}^{-1}\text{s}^{-1}\text{keV}^{-1}$ ) for 6 MeV electrons



# Do you lose electrons near 90° ?



# Acknowledgements

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# Combined EMIC and Chorus diffusion rates

$L^* = 4.5$

